UNRAVELLING SPECTRAL SIGNATURES AND PHOTOCHEMICAL PROCESSES OF PAHS

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The Physics and Chemistry of the Interstellar Medium Celebrating the first 40 years of Xander Tielens' contribution to Science Palais des Papes, Avignon, 3 Sep. 2019

OUTLINE

- The physics and chemistry of the ISM IS PAHs
- Electronic signatures of IS PAHs
- Photochemistry of PAHs as driver for D storage

THE PHYSICS AND CHEMISTRY OF IS PAHS

Cosmic Inventory?

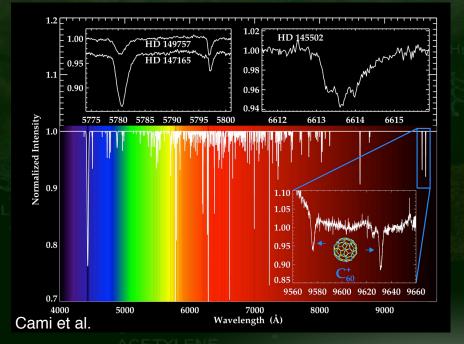


Isomerisation

Fragmentation

Clustering

THE PHYSICS OF IS PAHS How to identify IS PAHs?



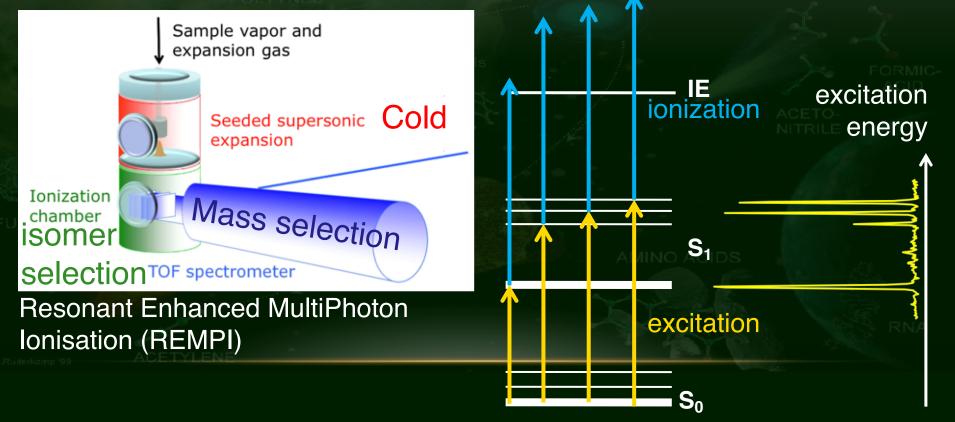
Limited correlation

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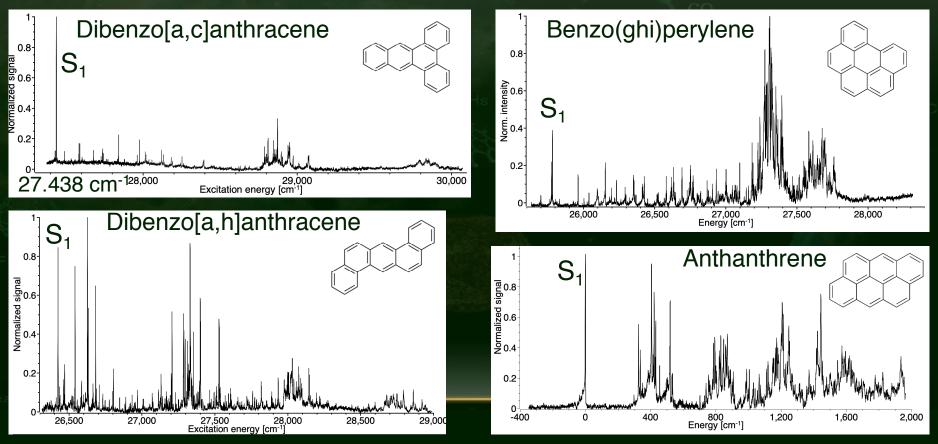
- Fortitude of possible isomers
- Individual abundances too low?
- No match found (yet)
- But... lack of accurate validated predictions and accurate experiments under IS conditions

Diffuse Interstellar Bands PAHs as carriers?

HIGH-RESOLUTION ACTION SPECTROSCOPY UNDER INTERSTELLAR CONDITIONS



GAS-PHASE ELECTRONIC SPECTRA

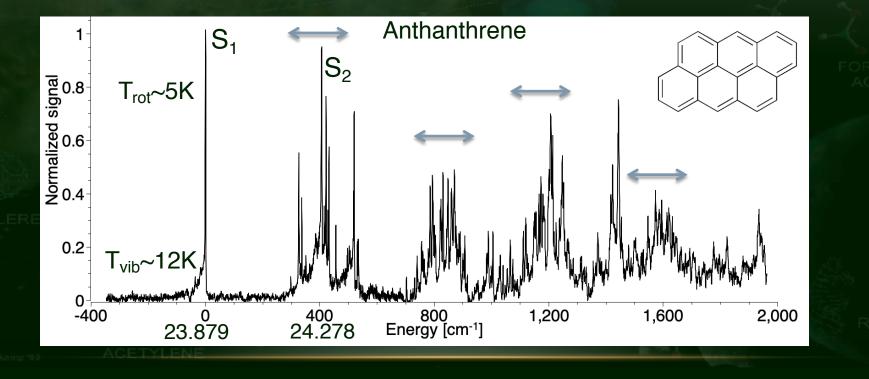


Roeterdink et al., *in preparation*

More results on poster of Hernán Velásquez Navarro

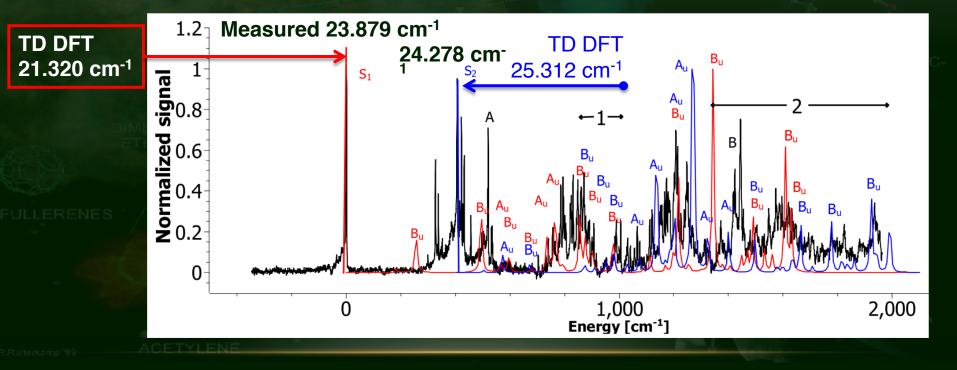
GAS-PHASE ELECTRONIC SPECTRA

ETHANE



Roeterdink et al., in preparation

GAS-PHASE ELECTRONIC SPECTRA



~11% deviation of $S_0 \rightarrow S_1$ transitions

Roeterdink et al., in preparation

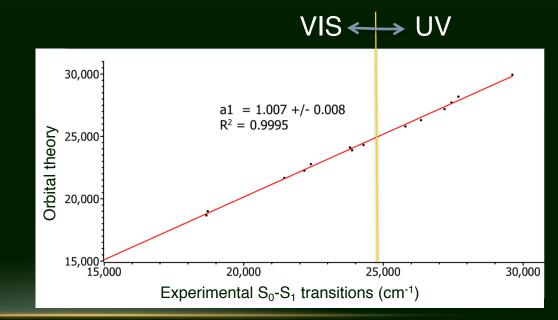
VALIDATE AND IMPROVE PREDICTIONS

- Cold high-resolution laboratory data of isolated PAHs
- TD DFT <15%

Roeterdink et al., in preparation (2)

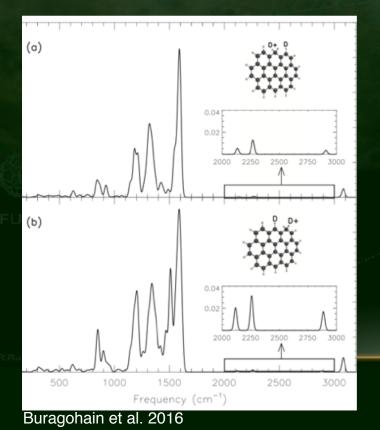
VALIDATE AND IMPROVE PREDICTIONS

- Cold high-resolution laboratory data of isolated PAHs
- TD DFT <15%
- Predictions ~1%!



Roeterdink et al., in preparation (2)

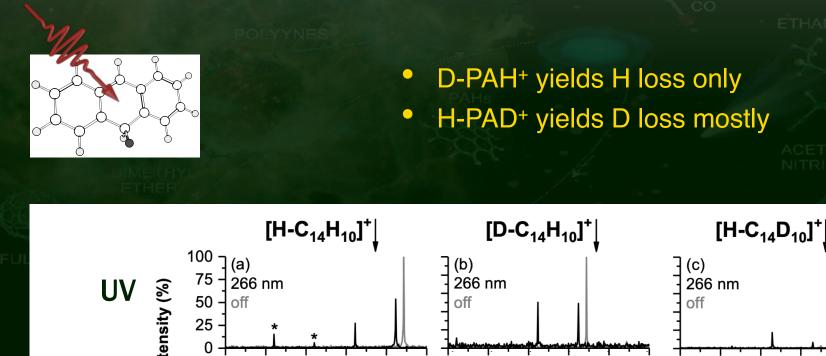
THE PHOTOCHEMISTRY OF IS PAHS PAHS AS SINK FOR DEUTERIUM



D/H ratio Primordial ~26 ppm ISM ~7 to 22 ppm
Predictions & observations
Experimental data lacking
Mechanism?

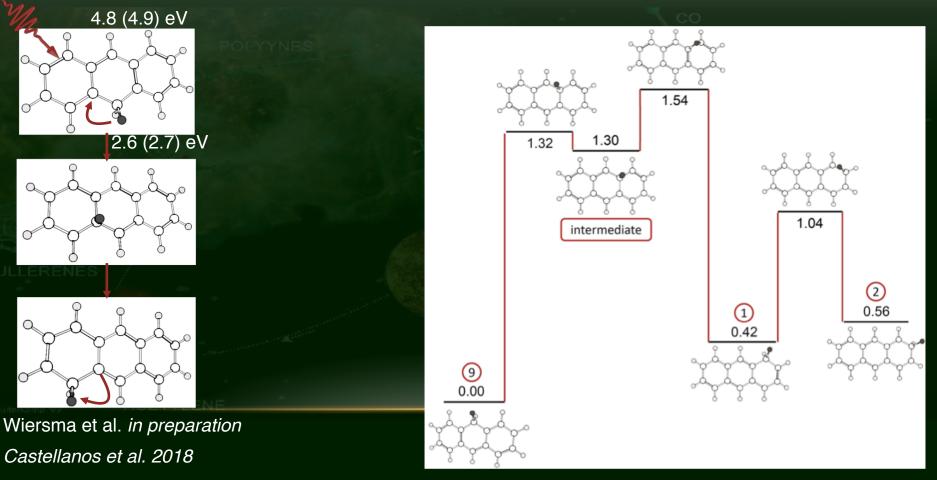
Could PAHs form a sink Does photochemistry play a role

PHOTOLYSIS OF D+-ANTHRACENE IONS



Wiersma et al. in preparation

PHOTO-INDUCED SCRAMBLING



CONCLUSIONS

Photochemistry of IS PAHs

- Photolysis induces scrambling in D-PAH⁺
- Driver for D storage
- Structure and size dependency?
- Re-interpretation of C-D band observations?

Electronic Signatures of IS PAHs

- Laboratory data under interstellar conditions
- Validated predictions within 1% accuracy

 Measuring DIB candidates in the lab!

ACKNOWLEDGEMENTS



Sandra Wiersma Wim Roeterdink Wybren Jan Buma

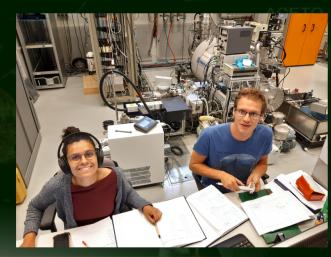
Sterrewacht Leiden

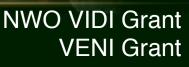
Infrared eXperiments

Alessandra Candian

Joost Bakker Giel Berden Jonathan Martens Jos Oomens

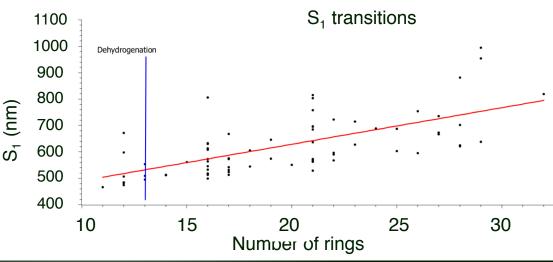
Thank you Xander





VALIDATE AND IMPROVE PREDICTIONS

- Cold high-resolution laboratory data of isolated PAHs
- TD DFT <15%
- Predictions <1%!
- Structure!



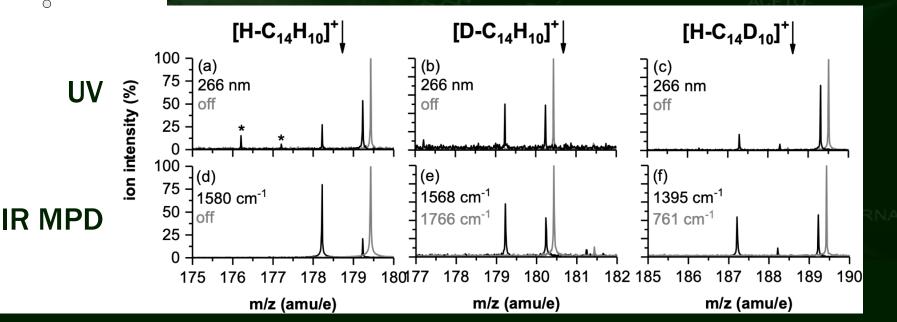
Roeterdink et al., in preparation (2)

PHOTOLYSIS OF D+-ANTHRACENE IONS

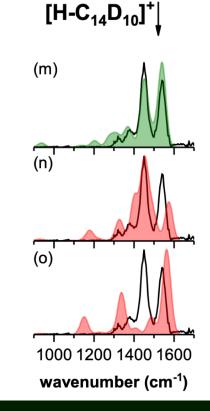
- D-PAH⁺ yields H loss only
- H-PAD⁺ yields D loss mostly
- IRMPD follows UV fragmentation

Wiersma et al. in preparation

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PHOTOLYSIS OF ANTHRACENE



 $[D-C_{14}H_{10}]^{+}]$ (j) (k) **(I)** 1000 1200 1400 1600 800 wavenumber (cm⁻¹)

[H-C₁₄H₁₀]⁺] (g) 70. 9 0.00 eV 6 (h) 0.42 eV (1) 2 0.56 eV 1000 1200 1400 1600 800 wavenumber (cm⁻¹)